Phys 301 Modern Physics

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Class 7: Energy/Momentum Conservation, 4-Vectors, **Causality**

Relativistic Force: Handout Part II

Consequence of Lorentz factor: nothing can be accelerated past speed of light.

A New Definition for Energy

1. At low velocity, the value *E* of the new definition should match the classical definition.

$$
W_{12} = \int_{1}^{2} \vec{F} \cdot d\vec{s} = K_2 - K_1
$$

$$
K = \frac{p^2}{2m}
$$

2. The total energy (Σ*E*) of an isolated system of bodies should be conserved in all inertial frames.

We will not go through the derivations but will practice applications.

Total Energy

There is energy associated *with mass itself:* rest energy E_0 Rest energy $E_0 = mc^2$ is an INVARIANT quantity.

$$
E = K + mc^2 = \gamma_p mc^2
$$

This definition of the relativistic *mass-energy E* fulfills the condition of conservation of total energy. (Not proven here)

Kinetic Energy

$$
E = K + mc^2 = \gamma_p mc^2
$$

The relativistic kinetic energy *K* of a particle with a rest mass *m* is:

$$
K=\gamma_pmc^2-mc^2=(\gamma_p-1)mc^2
$$

 $K =$ $\frac{1}{\sqrt{2}}$ \mathbf{Z} mu^2 Note: This is very different from the classical

Relativistic Kinetic Energy: Handout Part III

For slow velocities, the relativistic energy equation gives the same value as the classical equation! Remember the binomial approximation for $\gamma_p\colon \gamma_p\approx 1+1$ $\frac{1}{1}$ \mathbf{Z} u^2 c^2

$$
\rightarrow K = \gamma_p mc^2 - mc^2
$$

$$
\approx mc^2 + \frac{1}{2}mc^2 \frac{u^2}{c^2} - mc^2
$$

$$
\approx \frac{1}{2}mv^2
$$

Which graph best represents the total energy of a particle (particle's mass $m > 0$) as a function of its velocity *u*, in special relativity? 0 *c u* 0 *c u* 0 *c u* 0 *c u EEEE* $E = \gamma_p mc^2 = K + mc^2$ b) d) a) c) Discuss: What should *E* be when $u = 0$? 2. What should E be when $u = \text{very}$ very fast? Another way to solve: What is the shape of γ_p as a function of speed?

• What is the rest energy of a 100 g ball?
\n
$$
E_0 = mc^2 = (0.1 \text{ kg}) (3.0 \times 10^8 \frac{\text{m}}{\text{s}})^2
$$

\n $= 9.0 \times 10^{15} \text{J}$

• What is its kinetic energy if moving at 3 m/s?

$$
K = \frac{1}{2}mu^2 = (0.1 \text{ kg}) \left(3 \frac{\text{m}}{\text{s}}\right)^2
$$

= 0.45 J

Wait... if rest energy is SO BIG... How did we not mess up Physics 211?

Equivalence of Mass and Energy

$$
E_{\text{tot}} = \sqrt{\frac{m}{m} m c^2} = 2mc^2
$$

Violates Conservation of Energy!

$$
E_{\text{tot},2} = Mc^2 = 2K + 2mc^2 = E_{\text{tot},1}
$$

We find that the total mass *M* of the final system is bigger than the sum of the masses of the two parts! $M > 2m$. Potential energy inside an object contributes to its mass!!!

How does nuclear power work?

- Nuclei: neutrons and protons
- Strong force
- The potential energy associated with the force keeping them together = binding energy $E_{\rm B}$
- Total rest energy of particle equals the sum of the rest energy of all constituents minus the total binding energy $E_{\rm B}$:

$$
\boxed{Mc^2 = \mathbf{\Sigma}\bigl(\,m_i\ c^2) - E_B}
$$

Mass per nucleon

Important Relation

 $E = \gamma_p mc^2$ Total energy of an object:

Relativistic momentum of an object: $\vec{p} = \gamma_p m \vec{u}$

Energy – momentum relation: $E^2 = (pc)^2 + (mc^2)^2$

Momentum of a massless particle: *p =E/c* Velocity of a massless particle: $u = c$

$$
\vec{R} = \begin{bmatrix} ct \\ x \\ y \\ z \end{bmatrix} = \begin{bmatrix} ct \\ \vec{r} \end{bmatrix} \qquad \qquad \vec{P} = \begin{bmatrix} E \\ p_x c \\ p_y c \\ p_z c \end{bmatrix} = \begin{bmatrix} E \\ \vec{p}c \end{bmatrix}
$$

- Spacetime 4-vector
- Relationship defines one invariant quantity: spacetime interval.

• Energy-Momentum 4-vector

• Relationship defines other invariant quantity: rest mass energy.

Energy – momentum relation:

$$
(c\Delta t)^{2} - (\Delta x)^{2} = (\Delta s)^{2} \qquad E^{2} - (pc)^{2} = (mc^{2})^{2}
$$

$$
(c\Delta t)^2 - (\Delta x)^2 = (\Delta s)^2
$$

The Lorentz Transformation is a Matrix

$$
\begin{bmatrix} ct' \\ x' \\ y' \\ z' \end{bmatrix} = \begin{bmatrix} \gamma & -\beta\gamma & 0 & 0 \\ -\beta\gamma & \gamma & 0 & 0 \\ 0 & 0 & 1 & 0 \\ 0 & 0 & 0 & 1 \end{bmatrix} \begin{bmatrix} ct \\ x \\ y \\ z \end{bmatrix}
$$

$$
ct' = \gamma(ct - \beta x) \qquad t' = \gamma(t - v/c^2 x)
$$

$$
x' = \gamma(-\beta ct + x) \qquad x' = \gamma(x - vt)
$$

Conservation of 4-Momentum: Handout

What is causality?

- •**Causally connected** events *must* happen in a certain order in time.
- •Event being caused must follow event that causes it.
- •*…in all reference frames.*

Here is an event in spacetime.

Any light signal that passes through this event has the dashed world lines. These identify the '*light cone'* of this event.

Now we have two events A and B as shown on the left.

The space-time interval $(\Delta s)^2$ of these two events is:

A) Positive B) Negative C) Zero

Imagine you are in a rocketship moving past this reference frame. According to you, the space-time interval $(\Delta s)^2$ of these two events is:

Positive Negative C) Zero D) Cannot Be Determined

 $(\Delta s)^2$ is invariant under Lorentz transformation.

Which statement is true?

- A) Event A could cause Event B.
- B) Event B could cause Event A.
- C) Both answers above.
- D) Neither answer above.

For A to cause B, "information" would have to travel faster than the speed of light.

1) For each shaded region: What is the value of $(\Delta s)^2$? 2) Does each shaded region contain events *that can cause A*, events *that can be caused by A,* or events that are *causally unconnected?*

 $(\Delta s)^2 > 0$: **Time-like events** $(A \rightarrow D)$ There exists a frame where the two events could happen in the same place

 $(\Delta s)^2$ <0: **Space-like events** $(A \rightarrow B)$ There exists a frame where the two events could be simultaneous.

 $(\Delta s)^2$ =0: **Light-like events** $(A \rightarrow C)$